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# **A 10,000 yr record of high-resolution Paleosecular Variation from a flowstone of Rio Martino Cave, Northwestern Alps, Italy**

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## **Abstract**

Speleothems are potentially excellent archives of the Earth's magnetic field, capable of recording its past variations. Their characteristics, such as the continuity of the record, the possibility to be easily dated, the almost instantaneous remanence acquisition and the high time -resolution make them potentially unique high-quality Paleosecular Variation (PSV) recorders. Nevertheless, speleothems are commonly characterized by low magnetic intensities, which often limits their resolution. Here we present a paleomagnetic study performed on two cores from a flowstone from the Rio Martino Cave (Western Alps, Italy). U/Th dating indicates that the flowstone's

25 deposition covers almost the entire Holocene, spanning the period ca. 0.5-9.0 ka, while an  
26 estimation of its mean growth rate is around 1 mm per 15 years. The flowstone is composed of  
27 columnar calcite, characterized by a highly magnetic detrital content from meta-ophiolites in the  
28 cave's catchment. This favourable geological context results in an intense magnetic signal that  
29 permits the preparation and measurement of thin(~3 mm depth equivalent) samples, each  
30 representing around 45 yr. The Characteristic Remanent Magnetization (ChRM), isolated after  
31 systematic stepwise Alternating Field demagnetization, is well defined, with Maximum Angular  
32 Deviation (MAD) generally lower than 10°. Paleomagnetic directional data allow the  
33 reconstruction of the PSV path during the Holocene for the area. Comparison of the new data with  
34 archeomagnetic data from Italian archeological and volcanic records and using the predictions of  
35 the SHA.DIF.14k and pfm9k.1a global geomagnetic field models shows that the Rio Martino  
36 flowstone represents an excellent recorder of the Earth's magnetic field during the last 9,000  
37 years. Our high resolution paleomagnetic record, anchored by a high-quality chronology, provide  
38 promising data both for the detection of short term geomagnetic field variations and for  
39 complementing existing regional PSV curves for the prehistoric period, for which well-dated data  
40 are still scarce.

41

42 **Keywords:** Paleosecular variation, Rock magnetism, Speleothem, Italy

43

## 44 **1. Introduction**

45 To investigate geomagnetic field behavior in the past and to explore its short-term features, high-  
46 resolution records from globally distributed archives of different origin are necessary (Mandea and  
47 Olson, 2009). For Paleosecular Variation (PSV) reconstructions, an ideal paleomagnetic record

48 should satisfy several requirements, such as having a stable remanent magnetization, being well  
49 dated, offering a continuous record and presenting high-time resolution. Even though some Earth  
50 materials may satisfy a number of these, the characteristics of continuity and high-resolution are  
51 rarely coupled. Marine and lacustrine sediment sequences are best at ensuring continuous records  
52 and have therefore been intensively studied to obtain geomagnetic data over long time scales (e.g.  
53 Turner and Thompson, 1981; Rolph et al., 2004; Vigliotti, 2006). However, sometimes data  
54 reliability may be questionable: the remanence acquisition mechanisms, the smoothing effects of  
55 bioturbation, the inclination error and the remanence acquisition delay are just some of the  
56 problems that may affect this kind of record. On the other hand, volcanic rocks and fired  
57 archeological artifacts may preserve very reliable paleomagnetic data but they are highly  
58 discontinuous in time. The age uncertainties of the volcanic products, as well as the lack of  
59 continuity and the limited time extension of available *in situ* archeological baked clay structures,  
60 restrict their use for high-resolution record studies.

61 Several research groups have studied speleothems for both PSV and paleoenvironmental  
62 reconstructions (e.g. Latham et al., 1989; Lean et al., 1995; Openshaw et al., 1997; Osete et al.,  
63 2012; Xie et al., 2013; Font et al., 2014; Jaqueto et al., 2016; Lascu et al., 2016), revealing their  
64 high potential for magnetic and secular variation reconstructions (Lascu and Feinberg, 2011).  
65 Paleomagnetic time series from speleothems, although still sparse, can provide excellent temporal  
66 resolution if the speleothem has grown continuously and over a considerable age range, as, for  
67 example, in the case of the Mexican stalagmite studied in the pioneering work of Latham et al.  
68 (1986). The key features of speleothems are that they can grow continuously for  $10^3$ - $10^5$  yr and  
69 can be accurately dated by the uranium-series method (e.g. Richards and Dorale, 2003). They  
70 normally show little or no secondary alteration, and are generally easy to orient and sample  
71 (though with obvious consideration of natural heritage values).

72 Based on magnetic properties, the remanent magnetization of speleothems can be divided into  
73 two main genetic forms, detrital (DRM) or chemical (CRM) (Lascu and Feinberg, 2011). Detrital  
74 input can originate from flood or drip water sources (Openshaw et al., 1997; Fairchild et al., 2006).  
75 Moreover, speleothems present the advantage of acquiring their magnetization rapidly after  
76 formation, meaning that the registered magnetic remanence variations reliably reflect the PSV  
77 path in the past. Nevertheless, these promising features are confounded by a speleothem's  
78 generally low concentration of magnetic minerals, and thus their low magnetic signal, which limits  
79 their use in magnetic studies. To bypass this problem, large samples have been commonly used in  
80 paleomagnetic studies, but this reduces the time-resolution of the sample. Generally, a sample of  
81 around 2 cm may average ca 100-4000 yr (Strauss et al., 2013) and thus the obtained SV time-  
82 resolution is very low.

83 This paper reports the results of a paleomagnetic study performed on a flowstone sampled at Rio  
84 Martino Cave (North Western Alps, Italy). The favourable geologic context of the cave, which is  
85 mainly surrounded by meta-ophiolites, makes this flowstone very rich in detrital ferromagnetic  
86 components, and thus an ideal geomagnetic field recorder due to its high magnetic remanence  
87 properties. Although a high content of detrital material can compromise U/Th dating (Hellstrom,  
88 2006), we have been able to produce a continuous, radiometrically-dated, directional SV record  
89 for the area during the last ~10 kyr, at a sampling resolution averaging 45 yr. Comparison of the  
90 new data with archeomagnetic data from Italian artifacts and volcanic rocks and using predictions  
91 of global geomagnetic field models, shows that the Rio Martino flowstone represents an excellent  
92 recorder of the Earth's magnetic field in the past and demonstrates the potential of speleothems  
93 for PSV studies and for the investigation of short-term variations of the geomagnetic field.

94

## 95 **2. Geological setting and sampling**

96 Rio Martino Cave (44°42' N, 7°09' E) is located in the inner sector of the Western Alps (Northern  
97 Italy), which consists of a range of continental and oceanic tectono-metamorphic units bounded  
98 by major orogen-scale faulting (Balestro et al., 2014), and exhumed and stacked in the axial sector  
99 (Fig. 1). The cave is developed within the Mesozoic carbonate cover of the Palaeozoic Dora Maira  
100 (Balestro et al., 2013). This unit is overlain by the Monviso meta-ophiolite complex, a major  
101 eclogized remnant of the Ligurian-Piedmont oceanic lithosphere, which in turn is tectonically  
102 overlain by the Queyras Schistes Lustrés, interpreted as a fossil accretionary wedge.

103 The surface above the cave is overlain mainly by glacial deposits. The cave is located at 1530 m  
104 a.s.l. on the right flank of the upper Po valley. It is a spring cave, ca. 3000 m long, with 200 m of  
105 elevation difference, and is crossed by a small river with an average discharge of 50 l/s (maximum  
106 200 l/s) (Badino and Chiri, 2005).

107 The presence of highly magnetized rocks in the cave's catchment (Fig. 1) and the strong magnetic  
108 anomalies observed in the Monviso Massif area (Lanza and Meloni, 2006) could induce a magnetic  
109 deflection effect in the area. To evaluate the possible effect exerted by the meta-ophiolitic masses  
110 and to confirm that it does not exert a significant influence on the paleomagnetic sampling, we  
111 used a triaxial fluxgate magnetometer to measure the geomagnetic field components outside and  
112 next to the entrance, as well as inside the cave. The computed magnetic inclination values of 60.7°  
113 outside the cave and 60.5° on the flowstone surface are fully comparable to the 2013 IGRF model  
114 of 60.6° (<http://www.ngdc.noaa.gov/geomag-web>). Besides, outside the cave we performed some  
115 orientation checks by using both the magnetic and the solar compass. The difference between the  
116 two declinations was small, ranging from -5° to +2°. Such differences are insignificant and indicate  
117 that any local magnetic effects on the paleomagnetic sampling can be considered negligible.

118 Two sampling campaigns were carried out to collect two cores from the same flowstone, which  
119 has accumulated on the side of a seasonally active stream with a high-detrital content. The cores

120 were taken ~20-30 cm apart and drilled using an adapted electric-powered drill. The first core  
121 (RMD1), sampled during a campaign in 2010, was not azimuthally oriented. The second core  
122 (RMD8), sampled in 2013, was oriented *in situ* by magnetic compass and inclinometer (figure S1 in  
123 the supplemental material). Each core was ca. 60 cm long and was drilled perpendicular to the  
124 flowstone growth axis.

125 A quarter of each core was dedicated to paleomagnetic analysis. The investigated sub-samples  
126 consisted of small slices, about 3 mm thick (varying from 2.5 to 4 mm), cut almost perpendicular to  
127 the speleothem's growth direction. Slicing was performed using a very thin non-magnetic saw,  
128 which ensured that only 1 mm of material was consumed during the cut. Following this systematic  
129 sampling, we obtained 146 slices from RMD1 and 143 from RMD8. Each slice was positioned in the  
130 centre of a non-magnetic plastic cylinder (2.5 cm diameter, 2.3 cm height) that allowed its  
131 handling as per standard paleomagnetic samples (Fig. 2).

132

### 133 **3. Methods**

#### 134 *3.1. U/Th dating and age modelling*

135 Nineteen solid prisms of ~40 mg (~2 mm wide along the lamina and 1 mm thick on growth axis)  
136 from RMD1 were used for age determination (Table S1 in Supplementary Material). The U/Th  
137 dating was performed at the University of Melbourne (Victoria, Australia) following the method of  
138 Hellstrom (2003). Briefly, samples were dissolved and a mixed  $^{236}\text{U}$ - $^{233}\text{U}$ - $^{229}\text{Th}$  spike was added  
139 prior to removal of the carbonate matrix with ion-exchange resin. The purified U and Th fraction  
140 was introduced in a dilute nitric acid to a multi-collector inductively coupled plasma mass  
141 spectrometer (MC-ICPMS, Nu-Instruments Plasma). The  $^{230}\text{Th}/^{238}\text{U}$  and  $^{234}\text{U}/^{238}\text{U}$  activity ratios  
142 were calculated from the measured atomic ratios using an internally standardised parallel ion-

143 counter procedure and calibrated against the HU-1 secular equilibrium standard. Correction for  
144 detrital Th content was applied using initial activity ratios of detrital thorium ( $^{230}\text{Th}/^{232}\text{Th}$ )<sub>i</sub> of  $1.3 \pm$   
145 0.45. This value, and its relative  $2\sigma$  uncertainty, was calculated using a Monte Carlo 'stratigraphic  
146 constraint' procedure based on the series of U/Th ages (Hellstrom, 2006). A depth-age model was  
147 constructed using a Bayesian Monte Carlo approach following the method described by Drysdale  
148 et al. (2005) and Scholz et al. (2012).

### 149 *3.2. SEM-EDS analysis*

150 The mineralogy of the detrital inclusions in the studied flowstone was investigated by dissolving  
151 different portions of various thin slabs of the RMD1 core in diluted hydrochloric acid and passing  
152 the digests through 0.45 micrometre cellulose acetate filters. The residues, bearing almost all of  
153 the non-carbonate mineral inclusion types contained in the speleothem, were observed and  
154 analysed with a Cambridge Stereoscan 360 Scanning Electron Microscope housed at the Earth  
155 Science Department of the University of Torino, Italy. Analyses were performed using an Oxford  
156 Inca X-Act 200 EDS microanalysis equipped with a Link Pentafet detector (thin window), allowing  
157 qualitative/quantitative determination of light elements (down to boron). All data were obtained  
158 at 15 kV HT, 25 mm WD, probe current range 800 pA – 1.2 nA and analysis time from 60 to 500 s.  
159 Primary standardization was performed on SPI Supplies and Polaron Equipment standards, and the  
160 system was regularly calibrated against a high-purity metallic Co standard before each  
161 experimental session. Data were processed with the Inca 200 Microanalysis Suite Software,  
162 version 4.08, and calibrated on natural mineral standards using the ZAF correction method.  
163 Analytical data are considered to be only semi quantitative due to the nature of the samples  
164 (rough surface of the particles, lack of horizontality, lack of surface polishing). A total of about  
165 1500 analyses was performed on seven samples coming from different portions of the core,



166 corresponding to about 200 measurements for each filter, randomly scattered on the filter surface  
167 for better representativeness.

168 Despite the results of magnetic analysis, very few magnetite particles were found in the filtered  
169 material, most likely because the single-domain magnetic particles were not retained by the 0.45  
170  $\mu\text{m}$  filter. Magnetite was indeed observed in sandy materials from Rio Martino, being found in the  
171 bed sediments of the relatively high-energy environment of the cave stream, rather than as  
172 detritus in carbonate flowstone speleothems.

173

### 174 3.3. Rock magnetic measurements

175

176 All magnetic measurements were performed at the ALP Paleomagnetic Laboratory (Peveragno,  
177 Italy). Rock magnetic experiments were performed on representative samples from both cores.  
178 Rock magnetism was investigated by low-field susceptibility ( $k_m$ ) and natural remanent  
179 magnetization ( $J_r$ ) measurements using a KLY3 kappabridge and a JR6 spinner magnetometer with  
180 a sensitivity of the order of  $10^{-8}$  SI and  $10^{-6}$  A/m, respectively. Susceptibility was measured at least  
181 five times *per* sample in order to calculate a mean value. Standard deviation is low and normally  
182 less than 5% for specimens associated with a susceptibility spike; uncertainty grows to 20-35% for  
183 the remaining specimens, characterized by negative (diamagnetic) susceptibility values. All  
184 samples were weighed to get the mass-normalized susceptibility ( $\chi$ ,  $\text{m}^3\text{kg}^{-1}$ ) and intensity ( $J$ ,  
185  $\text{Am}^2/\text{Kg}$ ). Their values are represented as a function of the core depth in figure S2 of the  
186 supplemental material.

187 Isothermal Remanent Magnetization (IRM) curves were obtained with an ASC pulse magnetizer,  
188 applying stepwise increasing fields up to 1 T. Thermal demagnetization of a three-axis composite  
189 IRM was also performed on representative samples (Lowrie, 1990). An IRM was imparted with an

190 ASC pulse magnetizer along the sample's three orthogonal axes, applying first a maximum 1.5 T,  
191 then a medium 0.3 T and finally a minimum 0.1 T magnetic field. Crossover plots of IRM curves and  
192 alternating field (AF) demagnetization of the saturation IRM (SIRM) were carried out to investigate  
193 the magnetic grain size (Symons and Cioppa, 2000).  
194 Finally, hysteresis cycles were obtained by a Vibrating Sample Magnetometer (VSM, LakeShore  
195 7410 - Maximum applied field  $B_{\max} = \pm 1$  T;  $H = \pm 10.000$  Oe /  $976.000$  A/m) at the Istituto  
196 Nazionale di Ricerca Metrologica (INRIM, Torino) and interpreted by RockMag Analyzer 1.0  
197 software (Leonhardt, 2006). All specimens were AF demagnetized stepwise up to 100 mT with a  
198 ASC-D 2000 equipment. Representative twin specimens were also stepwise thermally  
199 demagnetized with a Schonstedt TSD-1 furnace.

200

## 201 **4. Results**

### 202 *4.1. Chronology*

203 All the U/Th ages obtained from RMD1 were in stratigraphic order within the associated  
204 uncertainties, except for two samples that were consequently rejected as outliers (Table S1 in  
205 Supplementary Material). Macroscopic and thin-section analyses of core RMD1 shows no growth  
206 interruption along its length. Age modelling performed on RMD1 indicates that the flowstone  
207 grew continuously between  $0.56 \pm 0.06$  ka and  $9.7 \pm 1.6$  ka b2k (Fig. 3). The mean growth rate is  
208  $0.058$  mm/yr, which implies a mean time-resolution of ca. 60 yr (3 mm specimen + 1 mm cut) for  
209 the PSV record. The time averaged in each 3 mm slice sample is ca. 45 yr. The age of RMD8 was  
210 inferred by comparing clearly visible growth layers (Fig. 2a) between the two cores, associated  
211 with spikes in both the magnetic mass susceptibility and mass magnetization (figure S2 of the  
212 supplemental material).

## 213 4.2. EDS

214 The mineralogy of the detrital portion in the RMD1 core is in strong accord with the composition  
215 of the surrounding lithology. Apart of the calcareous formation in which the cave has developed,  
216 the main rocks in the area are prasinites, amphibolites and serpentines. Minerals were grouped by  
217 similar chemistry, with some simplifications: as stated above, analyses were only semi-  
218 quantitative. The main identified groups are: iron oxides (without magnetite, discriminated by  
219 morphological features), magnesium silicates (other than serpentine), serpentine group, white  
220 mica group, feldspar, tremolite-actinolite amphiboles, other amphiboles (mainly hornblende),  
221 epidote group, chlorite group, quartz and accessories. The main minerals (Fig. 4) are represented  
222 by iron oxides (not distinguishable by chemistry for the reason explained above), magnesium  
223 silicates and serpentine group minerals. Iron oxides are mostly irregular in shape as if they had  
224 undergone reworking from the stream or by feedwater (Perkins, 1996). In few cases, a framboidal  
225 shape suggests *in situ* growth.

226

## 227 4.3. Magnetic mineralogy

228 The mass magnetic susceptibility of the specimens strongly varies. It mostly shows a prevailing  
229 diamagnetic phase with small negative values (from  $-7$  to  $0 \times 10^{-9} \text{ m}^3 \text{ kg}^{-1}$  with a mean value of  $-4 \times$   
230  $10^{-9} \text{ m}^3 \text{ kg}^{-1}$ ), alternating with high positive spikes, up to  $970 \times 10^{-9} \text{ m}^3 \text{ kg}^{-1}$ , suggesting a very low  
231 concentration of magnetic minerals in these specimens. Calcite bulk susceptibility is  $-12.09 \mu\text{SI}$ ; its  
232 mass susceptibility is about  $-4.46 \times 10^{-9} \text{ m}^3 \text{ kg}^{-1}$  (Almqvist et al., 2010). Since the literature value for  
233 the susceptibility of calcite refers to single crystal, we can assume that the mass susceptibility for  
234 calcite in the speleothem is slightly higher, because of mineral porosity. Assuming a constant  
235 diamagnetic contribution mostly due to calcite, the relative variability of magnetic susceptibility is

236 indicative of variations of the concentration of magnetic minerals: a mean  $\chi$  value of  $-4 \times 10^{-9} \text{ m}^3$   
237  $\text{kg}^{-1}$  can be assumed to be representative of the “standard” content in magnetite, while high  
238 values represent for pulses of higher detrital input.

239 The natural magnetization intensity ( $J_r$ ) strongly varies from specimen to specimen, being on  
240 average around  $1\text{-}10 \times 10^{-6} \text{ Am}^2 \text{ kg}^{-1}$  with spikes up to  $80 \times 10^{-6} \text{ Am}^2 \text{ kg}^{-1}$ . The variations of these  
241 two bulk parameters are correlated; the computed correlation coefficients are  $r = 0.87$  and  $r =$   
242  $0.76$  for RMD1 and RMD8, respectively. This corroborates the hypothesis that their values are  
243 essentially controlled by changes in concentration of the magnetic oxide.

244 IRM acquisition curves from representative samples saturate at relatively low field (around 0.3 T),  
245 indicating the presence of low coercivity minerals (Fig. 5a). The IRM acquisition data were further  
246 analyzed applying the MAX UnMix software (Maxbauer et al., 2016) to six flowstone slices and to  
247 sand collected from the bed of the cave spring. For three specimens, i.e. RM101 and RM134  
248 flowstones and the sand, only one magnetic component was identified with Bh and DP values fully  
249 consistent with detrital magnetite (Bh = 1.46, DP = 0.40). In the remaining cases, two components  
250 were detected, one as above, and the other pointing to higher-coercivity magnetite. The  
251 computed S-ratio ( $S_{0.3 \text{ mT}}$ ) ranges from 0.95 to 1.00.

252 During the thermal demagnetization of the orthogonal IRM components (Lowrie, 1990), two  
253 typical behaviors were observed, which are independent of the magnetization intensity of the  
254 specimens. The first (e.g. sample RM68a), representing about the 80% of the measured  
255 specimens, suggests that the primary remanence is dominated by a soft magnetic carrier,  
256 demagnetized at ca 350-450 °C, which is interpreted as a titanomagnetite (Fig. 5b). The second  
257 (e.g. sample RM44a), in the remaining 20%, is characterized by a first drop in the magnetization  
258 intensity between 200 and 300 °C, which may be speculatively related to the existence of  
259 maghemite (Pan et al., 2000), even though this evidence is not sufficient to unambiguously

260 identify this magnetic phase (Zhu et al., 2012). The presence of (titano)magnetite of detrital origin  
261 is easily justified considering the geologic context of the cave and it is probably originated from  
262 the highly magnetic rocks of the surrounding area, mostly meta-ophiolites (Balestro et al., 2013).  
263 The occurrence of small serpentinite lithics was also detected. In those cases, the Median  
264 Destructive Field (MDF), which is normally stable and around 50-60 mT, drops to 5-25 mT.  
265 Deflections from  $MDF = 50$  mT occur only where both the magnetic susceptibility and the  
266 magnetization intensity values are high. To check for these variations, we performed the  
267 experiment of Symons and Cioppa (2000) on some selected specimens, which were characterized  
268 by MDF ranging from 15 to 60 mT. It consists of a crossover plot, where the %SIRM is plotted as a  
269 function of the applied field, and using a logarithmic scale (Fig. 5c). The results suggest that,  
270 except specimen SP212, which is characterized by MD (titano)magnetite, samples mainly contain  
271 SD to PSD (titano)magnetite grains.

272 The modified Lowrie-Fuller method (Johnson et al., 1975), which represents a valid first-order  
273 indicator of grain-size composition (Font et al., 2014), was applied to some specimens with MDF in  
274 the range 45-50 mT. Results always show L-type behaviors with the Anhysteretic Remanent  
275 Magnetization (ARM) dominating the Isothermal Remanent Magnetization (IRM) during AF  
276 demagnetization treatment, corroborating the occurrence of SD grains.

277 These results are confirmed by the hysteresis cycles (Figure S3 in the supplemental material)  
278 performed on 2 specimens, one characterized by high  $\chi$ , one by low  $\chi$ . After correction for the  
279 diamagnetic effect, the curves show the occurrence of a low-coercivity phase with saturation at  
280 0.3-0.5 T; in both cases, the ratios  $M_{rs}/M_s$  and  $B_{cr}/B_c$  fall next to the PSD field and SD + MD mixing  
281 curve in the Day plot (Dunlop, 2002).

## 282 **5. Paleomagnetic directions**

### 283 *5.1. Natural Remanent Magnetization and the Anisotropy of Remanent Magnetization*

284 Speleothems can potentially offer very useful records of PSV and the remanence acquisition  
285 mechanisms in speleothems have been previously studied in detail (e.g., Lascu and Feinberg, 2011;  
286 Strauss et al., 2013, and reference therein). In order to provide a reliable PSV record, the  
287 magnetization should be acquired and locked soon after the calcium carbonate film is deposited  
288 on a speleothem (almost instantaneously). Following Strauss et al. (2013), lock-time for a  
289 speleothem is sub-annual and the magnetization is a DRM. Synchronicity between crystallization  
290 and magnetization has been tested experimentally by synthetic stalagmite growth (Morinaga et  
291 al., 1989), confirming the short time-lapse in acquiring magnetization parallel to the ambient field  
292 direction.

293 To test if this requirement is encountered in Rio Martino flowstone and thus to check for its  
294 reliability as a PSV recorder, we measured both the Anisotropy of Magnetic Susceptibility (AMS)  
295 and the Anisotropy of Isothermal Remanent Magnetization (AIRM) on two selected sets of  
296 samples from RMD1 (azimuthally non-oriented core), each comprising a time interval of ca 1000  
297 yr. The first set comprised 16 samples (SP200 to SP260) from  $4.26 \pm 0.23$  to  $3.30 \pm 0.03$  ka, and the  
298 second set 14 samples (SP346 to SP397) from  $7.76 \pm 0.12$  to  $6.91 \pm 0.11$  ka. A difference of  $20^\circ$  in  
299 the mean magnetic ChRM inclination distinguished these two sets of specimens. The AMS was  
300 measured by a KLY-3 kappabridge. The results obtained on both sets of samples show a well-  
301 defined (confidence angles  $< 15^\circ$ ) mean minimum susceptibility axis,  $k_3$ , which is statistically  
302 vertical and perpendicular to the flowstone growth laminae (Fig. 6a). For the AIRM measurements,  
303 each specimen was first AF demagnetized using a tumbling 2G demagnetizer at 60 mT peak field  
304 and then given an isothermal remanent magnetization (IRM) with a steady field of 20 mT using an  
305 AGICO PUM-1 pulse magnet. After measurement with the spinner magnetometer, the sequence  
306 was repeated for a total of 12 different orientations of the IRM in order to calculate the anisotropy  
307 tensor. The experiment (Fig. 6b-c) shows that for both sets, the maximum IRM anisotropy axis  $I_1$  is

308 concordant or statistically indistinguishable from the mean ChRM direction, showing no relation  
309 with the speleothem growth laminae. This shows that ChRM direction is due to the statistical  
310 alignment of the magnetic particle and fully agrees with the conclusions of Zhu et al. (2012), who  
311 performed both AMS and AIRM on stalagmites. They found that the AMS was dominated by the  
312 calcite fabric, being the minimum susceptibility axis  $k_3$  aligned perpendicular to the stalagmite  
313 growth laminae, while the AIRM fabric showed the maximum remanence axis  $I_1$  almost parallel to  
314 the NRM direction. These data all point to a detrital origin of the magnetization, with the  
315 geomagnetic field control in the orientation of the ferromagnetic minerals.

316

## 317 *5.2. Characteristic Remanent Magnetization determination*

318 Demagnetization results are represented by intensity-decay curves and plotted in Zijderveld  
319 diagrams (Fig. 7). Most of the specimens are characterized by a small viscous remanent  
320 magnetization (VRM), which is easily removed at an AF field of 15-20 mT. The remaining  
321 demagnetization path is linear and points to the origin, indicating a stable remanent  
322 magnetization; this component has been interpreted as the Characteristic Remanent  
323 Magnetization (ChRM). The ChRM direction is mostly well defined and characterized by low MAD  
324 values (lower than  $8^\circ$  for the 91% of the studied samples). AF and thermal demagnetization results  
325 obtained from twin specimens are very similar (Fig. 7), confirming the reliability of the ChRM  
326 direction (Fig. 8; Table S2 in Supplementary Material). In the thermal demagnetization (Fig. 7b and  
327 7d), an inflection in the intensity decay curve it is observable at ca. 150 °C. As suggested by Strauss  
328 et al. (2013) the decay could be correlated with the occurrence of goethite, not detected however  
329 by any other experiment. The AF demagnetization treatment has been preferred rather than the  
330 thermal demagnetization, as it permits the further use of the same samples for  
331 paleoenvironmental and relative paleointensity investigations through Anhyseretic Remanent

332 Magnetization (ARM) measurements. Therefore, all samples were systematically AF demagnetized  
333 and ChRM directions were obtained from the AF demagnetization results. Demagnetization  
334 behavior in samples with low- and high-remanence (spike) does not change significantly except for  
335 specimens where serpentinite clasts were recognized.

336

## 337 **6. Directional Paleosecular Variation during Holocene**

338 Paleomagnetic directions obtained from the two cores (reported in Table S2 of Supplementary  
339 material) are plotted versus depth from the top of the core in Figure 9. Some spikes in declination  
340 show a strong correspondence with atypical MDF values, lower/higher than 20/60 mT. These  
341 deflected directions have been ascribed to the presence of small serpentinite lithic fragments and  
342 were thus rejected. The declination of RMD1, which was not azimuthally oriented, has been  
343 recovered after adjustment of its mean value to the Geocentric Axial Dipole (GAD) calculated at  
344 Rio Martino according to the following procedure: first, the mean direction of RMD1 has been  
345 calculated for the last 10 kyr and then its deviation from the GAD value has been computed. The  
346 difference in declination between the core and the GAD has been extracted from each declination  
347 value of the RMD1 core.

348 Generally, directions obtained from both RMD1 and RMD8 are in good agreement with each other  
349 and data reproducibility is high (Fig. 9). This is particularly evident for the inclination data at  
350 depths between 200.0 and 600.0 mm, where the two records match each other. Instead, in some  
351 cases, mostly at depths from 150.0 to 200.0 mm, differences in inclination of around 20°-25° are  
352 observed. The cause of such differences is not clear, even though uncertainties during sampling  
353 (slices not perfectly perpendicular to the flowstone growth) and deflections related to a possible  
354 anisotropy effect connected to the calcite crystals growth cannot be completely excluded. To



355 guarantee the high quality of the new data, only ChRM directions characterized by MAD values  
356 lower than  $6^\circ$  have been used in the plots. Directional data of each core have been kept clearly  
357 distinguishable, because only RMD1 was directly dated and even though the paleomagnetic  
358 records from the two cores are consistent and despite the fact that the two cores are only 30 cm  
359 apart, their simultaneous growth cannot be fully guaranteed.

360 The paleosecular variations registered by the Rio Martino speleothem are compared with spot  
361 archeomagnetic directions obtained from dated archeological structures and volcanic rocks from  
362 Italy. The data from the *in situ* archeological material are taken from the Italian archeomagnetic  
363 dataset (Tema et al., 2006; Tema, 2011) updated by some recently published results (Malfatti et  
364 al., 2011; Kapper et al., 2014; Tema et al., 2013; 2015; 2016). The data from the Italian volcanic  
365 eruptions have been downloaded from the GEOMAGIA50.v3 database (Brown et al., 2015). All  
366 data have been relocated at the geographic coordinates of Rio Martino via the virtual geomagnetic  
367 pole method (Noel and Batt, 1990). The comparison shows that the archeomagnetic and volcanic  
368 data generally fit very well to the speleothem directions (Fig. 10). Some discrepancies in  
369 declination can be observed around 1000 AD. For this period the speleothem declination values  
370 are very low and quite dispersed. Nevertheless, it is particularly interesting to note that for the BC  
371 period, the available archeomagnetic data, even if very limited and often accompanied by large  
372 error bars, are in excellent agreement with the new data, in spite of the speleothem inclination  
373 being systematically higher. Recently, Ponte et al. (2016) noticed that the inclination values in  
374 speleothems seem to vary as a function of the orientation of the calcite growth. To check for this  
375 effect, we calculated the mean paleomagnetic direction for both RMD1 and RMD8 applying Fisher  
376 Statistics, selecting ChRMs with  $MAD < 6^\circ$ . Their mean directions (RMD1:  $D = 358.8^\circ$ ,  $I = 65.1^\circ$ ,  $\alpha_{95}$   
377  $= 1.9^\circ$ ; RMD8:  $D = 2.4^\circ$ ,  $I = 65.8^\circ$ ,  $\alpha_{95} = 1.7^\circ$ ) are close to the GAD at Rio Martino's geographic  
378 coordinate ( $D = 0^\circ$ ,  $I = 63^\circ$ ), even if statistically distinguishable. However, AIRM shows a reliable

379 consistency between the mean ChRM and the mean  $I_1$  axis, which confirms that the orientation of  
380 the ferromagnetic minerals was mainly controlled by the geomagnetic field present during  
381 speleothem accumulation. This substantiates the high potential of Rio Martino speleothems to  
382 continuously and reliably register the Earth's magnetic field, offering a unique source of high  
383 quality data for the BC period where *in situ* archeological artifacts are very scarce.

384 The new data are also compared with the predictions of global geomagnetic field models. Here,  
385 we have used for comparison the pfm9k.1a (Nilsson et al., 2014) and the SHA.DIF.14k (Pavón-  
386 Carrasco et al., 2014) models that are the most recently published global geomagnetic models for  
387 the Holocene. There is good agreement between the speleothem records and the global models  
388 predictions, confirming some interesting features of the Earth's magnetic field in the past. The  
389 eastward declinations around 1000 BC mainly observed in the SHA.DIF.14k model are observed in  
390 the speleothem data for the same time period, and also show high declination values. For the  
391 4000-2000 BC period, only small declination variations are shown by the speleothem data, in  
392 agreement with the pfm9k model's predictions, while the declination peaks seen in the  
393 SHA.DIF.14k model (e.g. around 3600 BC) are not confirmed by the speleothem data. For periods  
394 older than 5000 BC, speleothem records show generally higher declination values compared to the  
395 model predictions and other archeomagnetic data. Interestingly, similar eastward declination  
396 values were found for cave sediments in Switzerland (Kapper et al., 2014). Regarding the  
397 inclination data, good agreement can be observed for the periods 6000-3500 BC and 500 BC-500  
398 AD. However, around 1000 BC, speleothems show an interesting high-inclination peak that is not  
399 observed in the models or sustained by the available archeomagnetic data. This peak is actually  
400 only observed on the data from the RMD8 core and definitely more independent records are  
401 necessary to investigate if it corresponds to a real abrupt directional change (as it corresponds also  
402 to high declination values) of the geomagnetic field at this time period. For the 7500 BC to 6000 BC

403 period, the speleothem records show continuously increasing inclination with a peak around 6000  
404 BC that seems to be in agreement with the model's predictions.

405

## 406 **7. Conclusions**

407 Some outstanding characteristics of the Rio Martino flowstone, such as its continuous growth, the  
408 well-constrained chronology and the intense magnetic signal, make its paleomagnetic directional  
409 record for the Holocene in the northwestern Italy particularly appropriate for PSV investigation.  
410 The high magnetic signal permits a high-resolution record of around 60 yr per data point; the  
411 regular scatter of paleomagnetic data through time shows an almost constant distribution of  
412 directional data through the Holocene.

413 The obtained directional results are well defined and offer a unique, almost continuous, secular  
414 variation record for the last ~10,000 years. Although some discrepancies can be observed,  
415 comparison with archeomagnetic data and global geomagnetic field models confirms the high  
416 potential of these speleothems to the reconstruction of the Earth's magnetic field variations in the  
417 past.

418 Our results show that the Rio Martino flowstones are not affected by recrystallization effects or  
419 secondary alterations. The speleothems do not show any inclination shallowing when compared  
420 with model predictions, and in some cases show high inclination peaks that are not observed by  
421 the models (e.g. around 3800 BC, 1000 BC, 800 AD).

422 The record characteristics overcome some typical features affecting both clastic sedimentary and  
423 the archeomagnetic PSV records, including the smoothness of the magnetic data in the case of the  
424 former and the presence of temporal gaps and uneven data distribution in the case of the latter.

425 The high resolution obtained points to the possibility of detecting short and abrupt geomagnetic  
426 field changes by studying a wide variety of Earth Magnetic Field variations at a timescale from tens

427 of years to the millennia and highlights the importance of regional differences when modelling the  
428 Earth's field. The use of speleothem records for PSV reconstructions can be particularly important  
429 for the prehistoric period where other sources of data coming from archeological artifacts or well-  
430 dated volcanic eruptions are scarce.

431

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436 coring operations. J.M. Feinberg and two anonymous referees are thanked for helpful reviews that  
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438 who enthusiastically supported this study during the last months of his life giving fundamental  
439 advices and suggestions.

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576

577 **Figure Caption**

578

579 Figure 1. a) Structural sketch map of the Western Alps; b) 3D reconstruction of the Inner Western  
580 Alps in the Rio Martino zone (modified after Balestro et al., 2014). The square indicates the  
581 location of the Rio Martino Cave. Note: the region highlighted in the small inset map is not the  
582 same as the one shown in the enlargement.

583 Figure 2. a) Part of core RMD8; b) part of the flowstone systematically cut and sampled in 3 mm-  
584 thick slices; c) the amagnetic plastic cylindrical holder created in order to fix the small samples in  
585 the centre of the cylinder and treat them as standard paleomagnetic samples.

586 Figure 3. Age-depth model for RMD1 core. The age is expressed both in ka AD and in b2k (before 2  
587 ka).

588 Figure 4. The distribution of the mineral species in the insoluble (detrital) fraction of Rio Martino  
589 speleothem. The picture is the sum of ca 1500 EDS determinations from seven different portions  
590 from the same core. The “accessory minerals” are species <2.5 % of the analyzed particles, for  
591 each sample and includes rutile, zircon, monazite, apatite (mainly apatite-F), sphene, xenotime,  
592 galena, pyrite, ilmenite, barite.

593 Figure 5. a) Isothermal remanent magnetization (IRM) acquisition curves; b) thermal  
594 demagnetization of a composite three-axes IRM (Lowrie, 1990); c) crossover plots (Symons and  
595 Cioppa, 2000).

596 Figure 6. Equal area stereographic projections of the eigenvectors for a) the anisotropy of  
597 magnetic susceptibility, and for b) and c) isothermal remanent magnetization, where the  
598 maximum, intermediate, and minimum eigenvectors are denoted by squares, triangles, and  
599 circles, respectively. The 95% confidence ellipses for the eigenvectors are shown by unfilled

600 ellipses. The mean ChRM directions and their  $\alpha_{95}$  errors for specimens SP200 to SP260 (b) and  
601 SP346 to SP397 (c) are denoted by stars with grey ellipses.

602 Figure 7. Thermal and AF demagnetization results from twin specimens from samples a-b) RM7  
603 and c-d) RM20 plotted in intensity decay plots (left) and Zijderveld diagrams (right). Symbols: full  
604 dots = declination; open dots = apparent inclination.

605 Figure 8. Equal-area projections of the ChRM directions for five samples obtained from a) AF and  
606 b) thermal demagnetization on twin specimens. The star represents the mean value calculated for  
607 each group of samples following a Fisherian distribution.

608 Figure 9. a) Declination and b) inclination data from cores RMD1 (red) and RMD8 (blue) plotted  
609 versus depth in mm from the top of the core.

610 Figure 10. a) Declination and b) inclination plots of the RMD1 (red) and RMD8 (blue) compared  
611 with the Italian archaeomagnetic data from archeological artefacts (green diamonds) and volcanic  
612 rocks (black squares) and the pfm9k (magenta line) and SHA.DIF.14k (black line) global  
613 geomagnetic field models. All directions are calculated at the geographic coordinates of Rio  
614 Martino ( $44.7^{\circ}$  N,  $7.15^{\circ}$  E). Age is given both as Calendar Age (year AD) and b2k (before 2 ka).

#### 615 **Figure caption of the Supplemental material**

616 Figure S1. a) Photograph of the flowstone's sampling and b) of the device to orient the core.

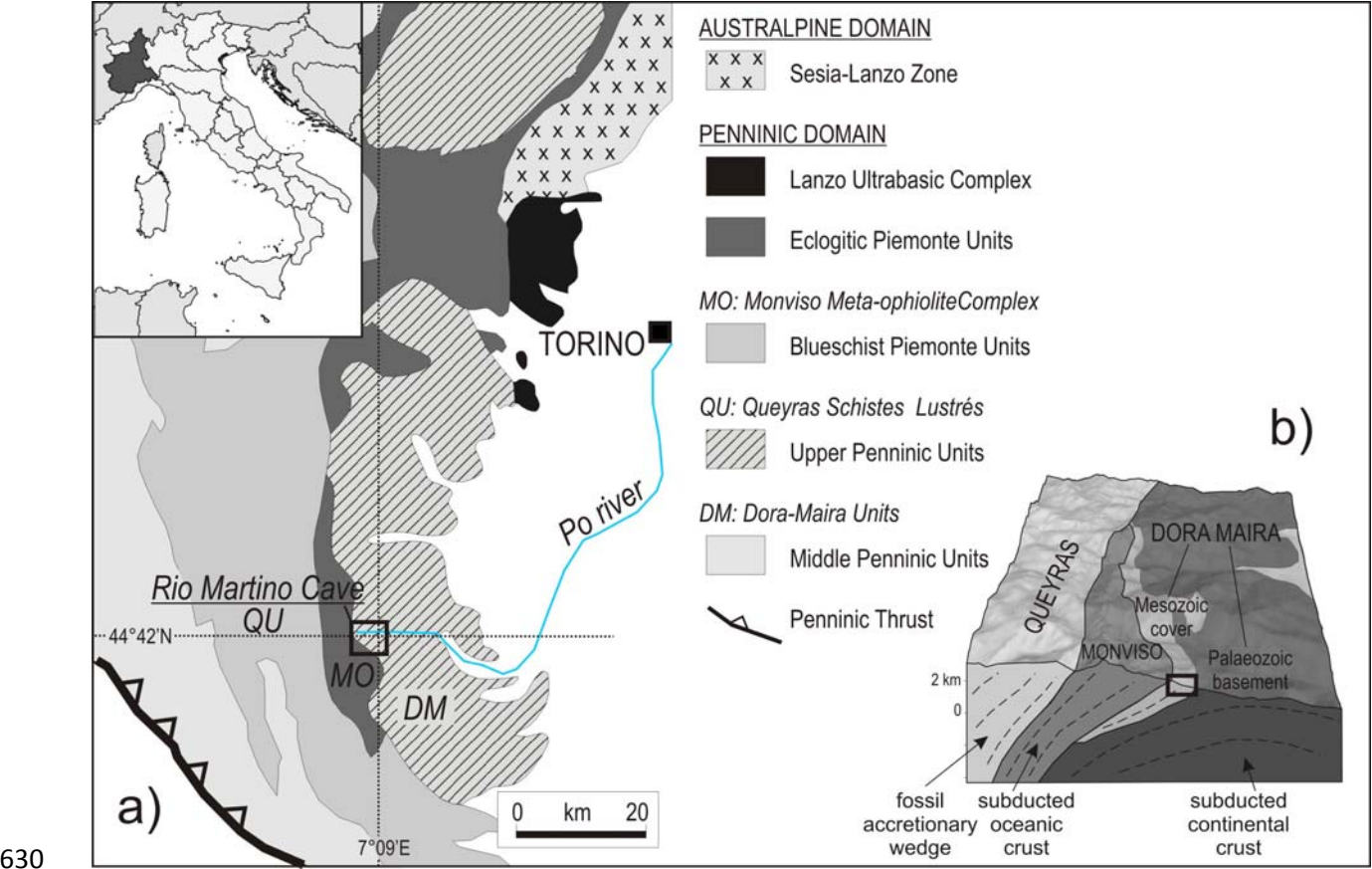
617 Figure S2. a) Mass susceptibility and b) mass remanence intensity from RMD1 (red) and RMD8  
618 (blue) plotted versus depth in mm from the top of the core.

619 Figure S3. Hysteresis curves (mass magnetization versus applied field) for specimens SP421 and  
620 SP460. In a) uncorrected curves; in b) slope-corrected for diamagnetic effect curves.

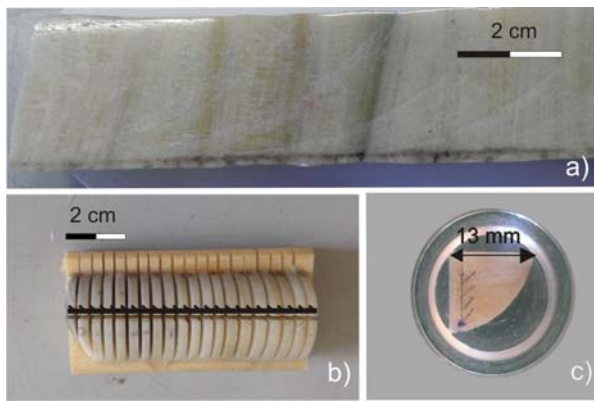
621 **Table caption**

622 Table S1. Corrected U/Th ages for RMD1 core. Isotope ratios are expressed as activity ratios  
623 standardized to the HU-1 secular equilibrium standard. Ages have been calculated using decay  
624 constants of  $9.195 \times 10^{-6}$  ( $^{230}\text{Th}$ ) and  $2.835 \times 10^{-6}$  ( $^{234}\text{U}$ ). Depths are from top.

625 Table S2. Characteristic remanent magnetization directions (ChRMs) of the samples from the  
626 RMD1 (left) and RMD8 (right) cores. Legend: z = depth in mm from the core top; D, I = magnetic  
627 declination and inclination; MAD = Mean Angular Deviation;  $D_{\text{corr}}$  = declination corrected by  
628 subtracting the angular difference between the  $D_{\text{GAD}} = 0^\circ$  and the mean ChRM declination ( $D =$   
629  $146.1^\circ$ ).

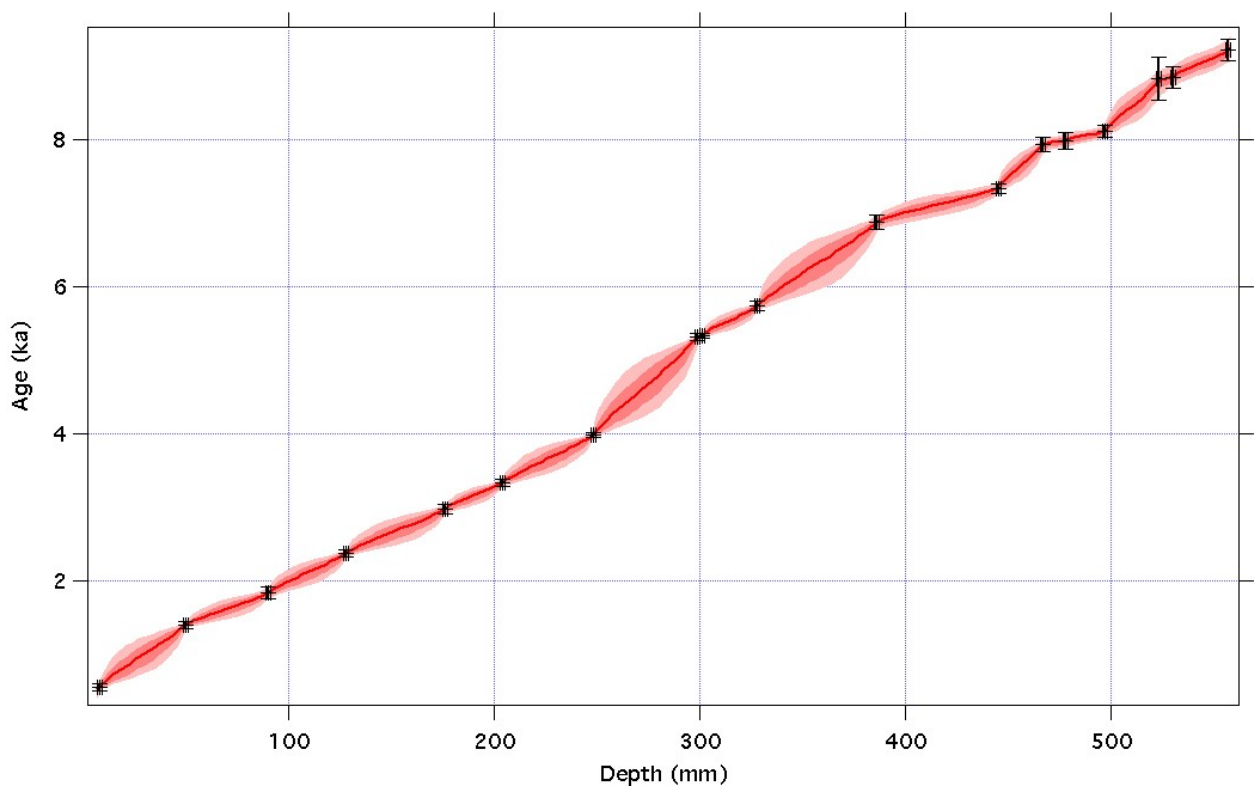


631 **Figure 1**



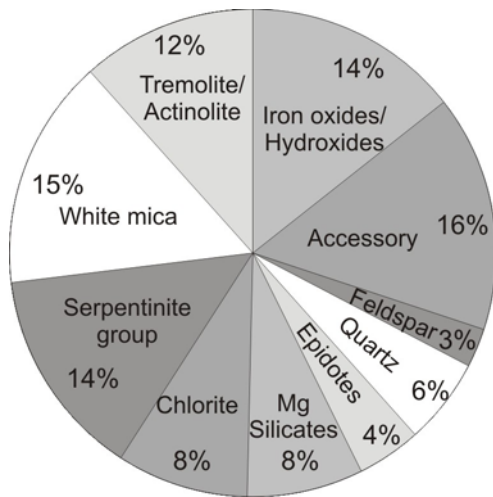
632

633 Figure 2



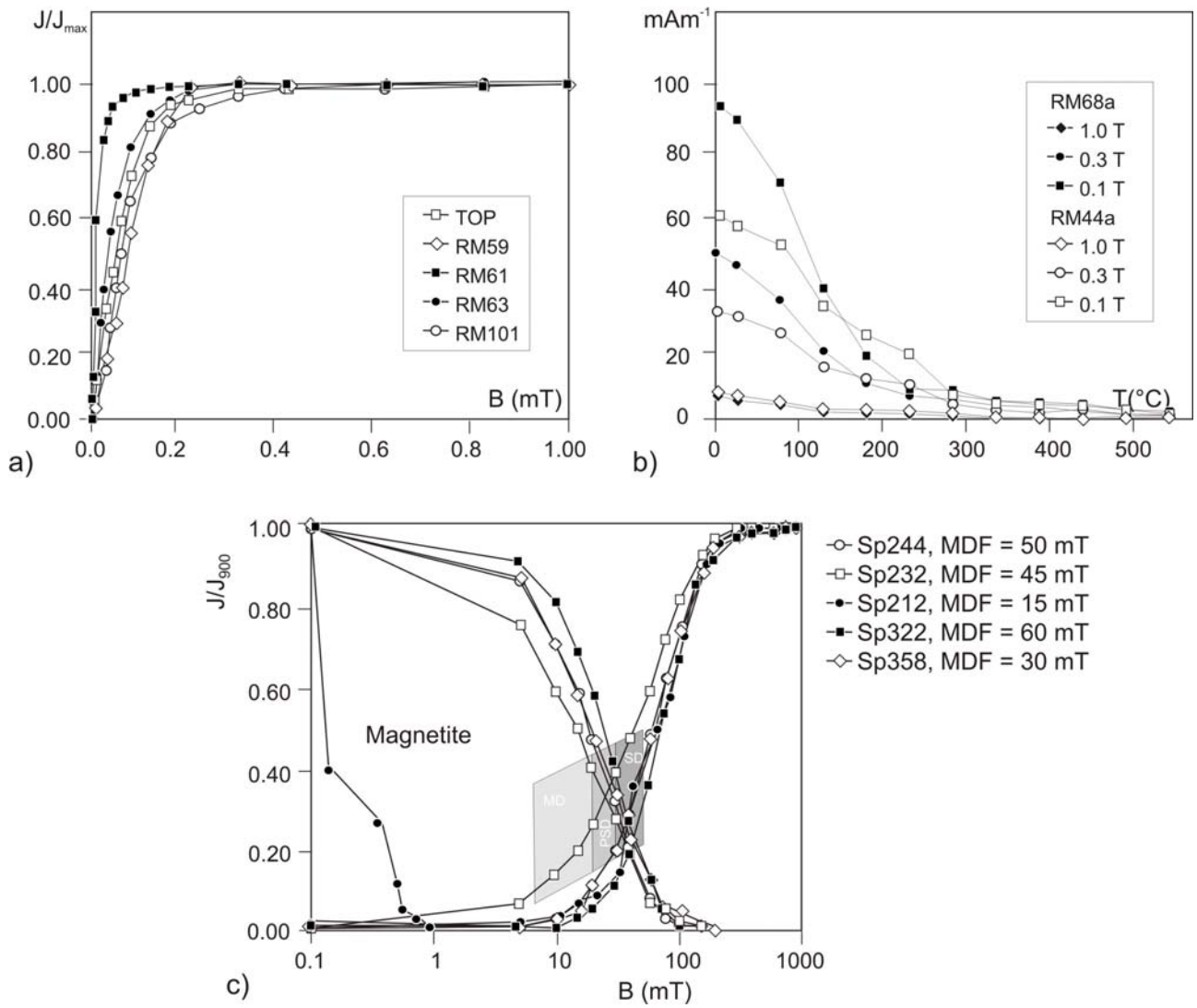
634

635 Figure 3



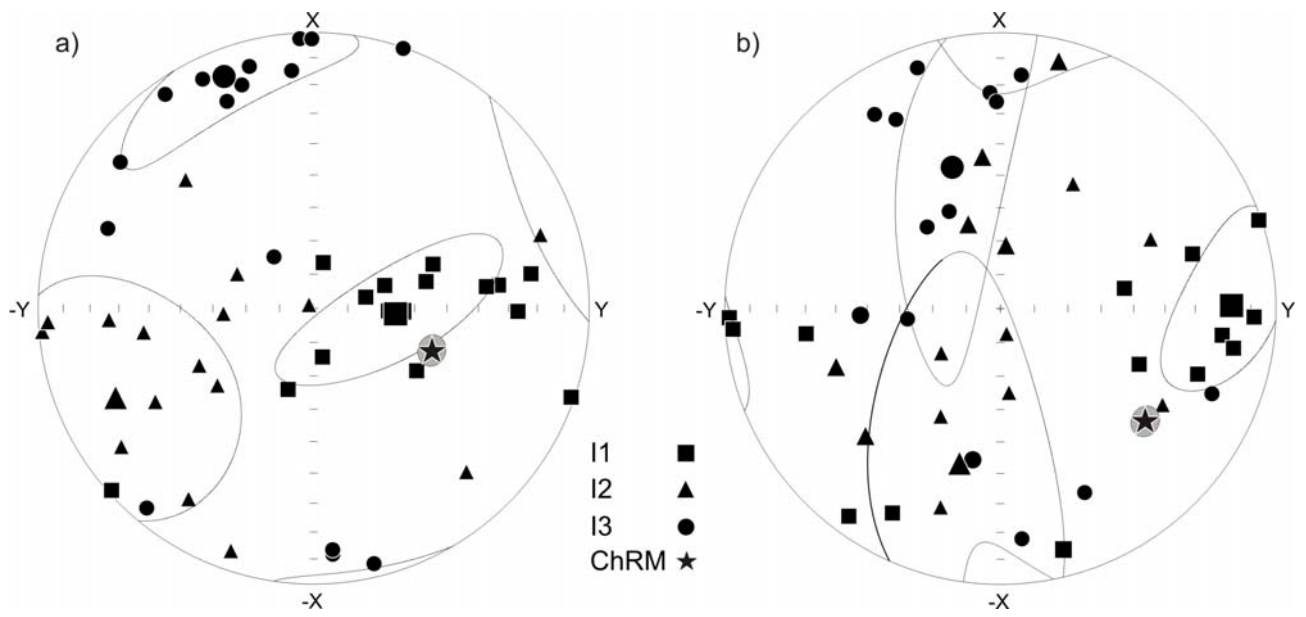
636

637 Figure 4



638

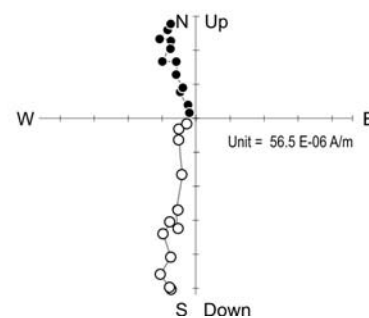
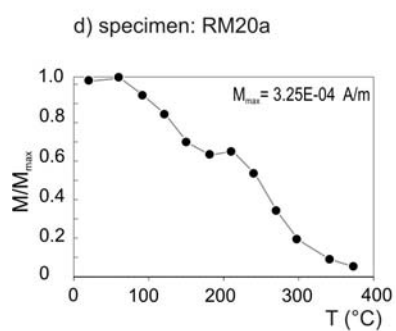
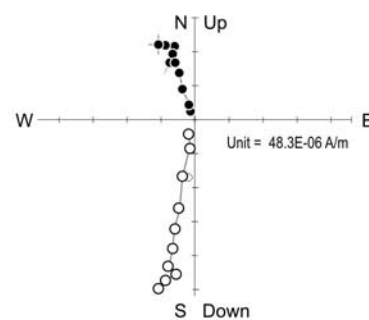
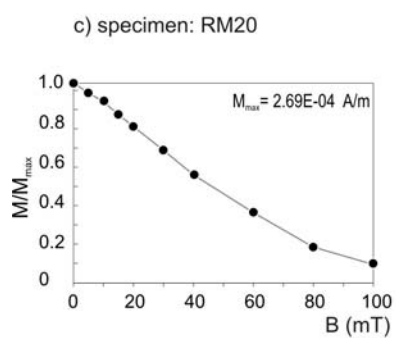
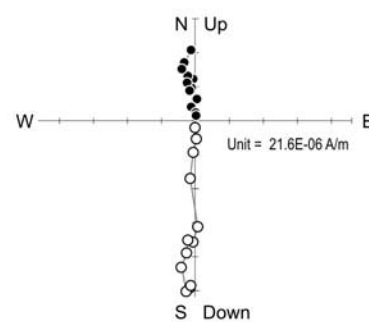
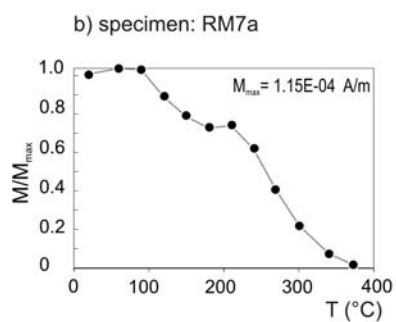
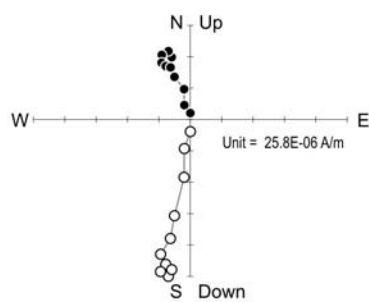
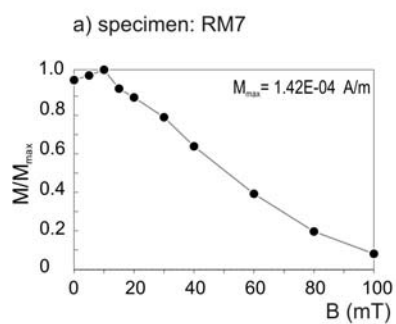
639 Figure 5



640

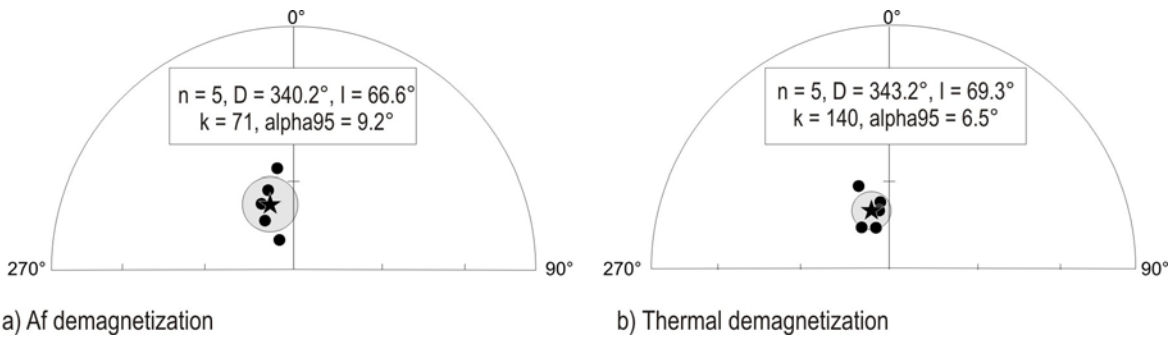
641 Figure 6



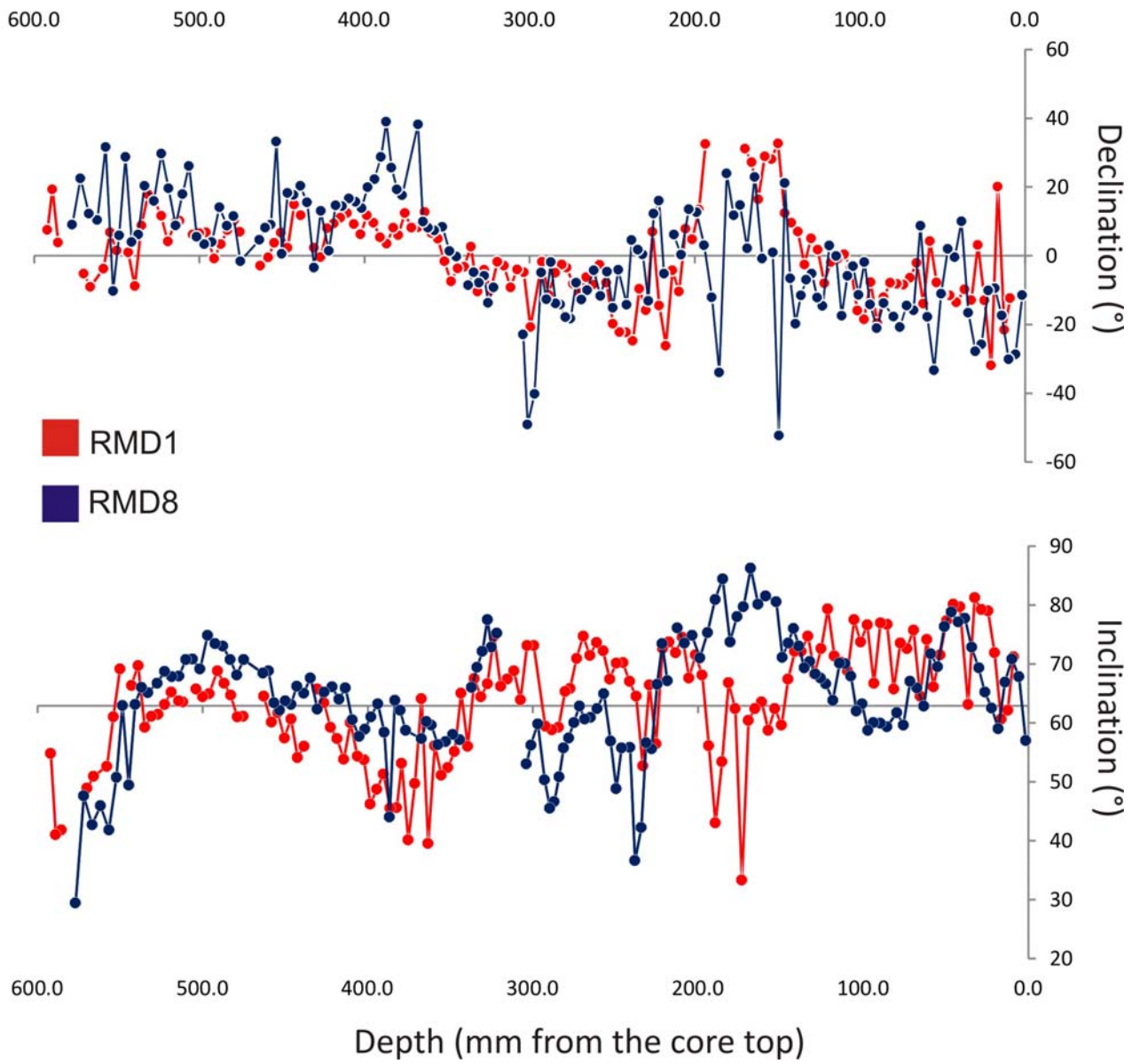


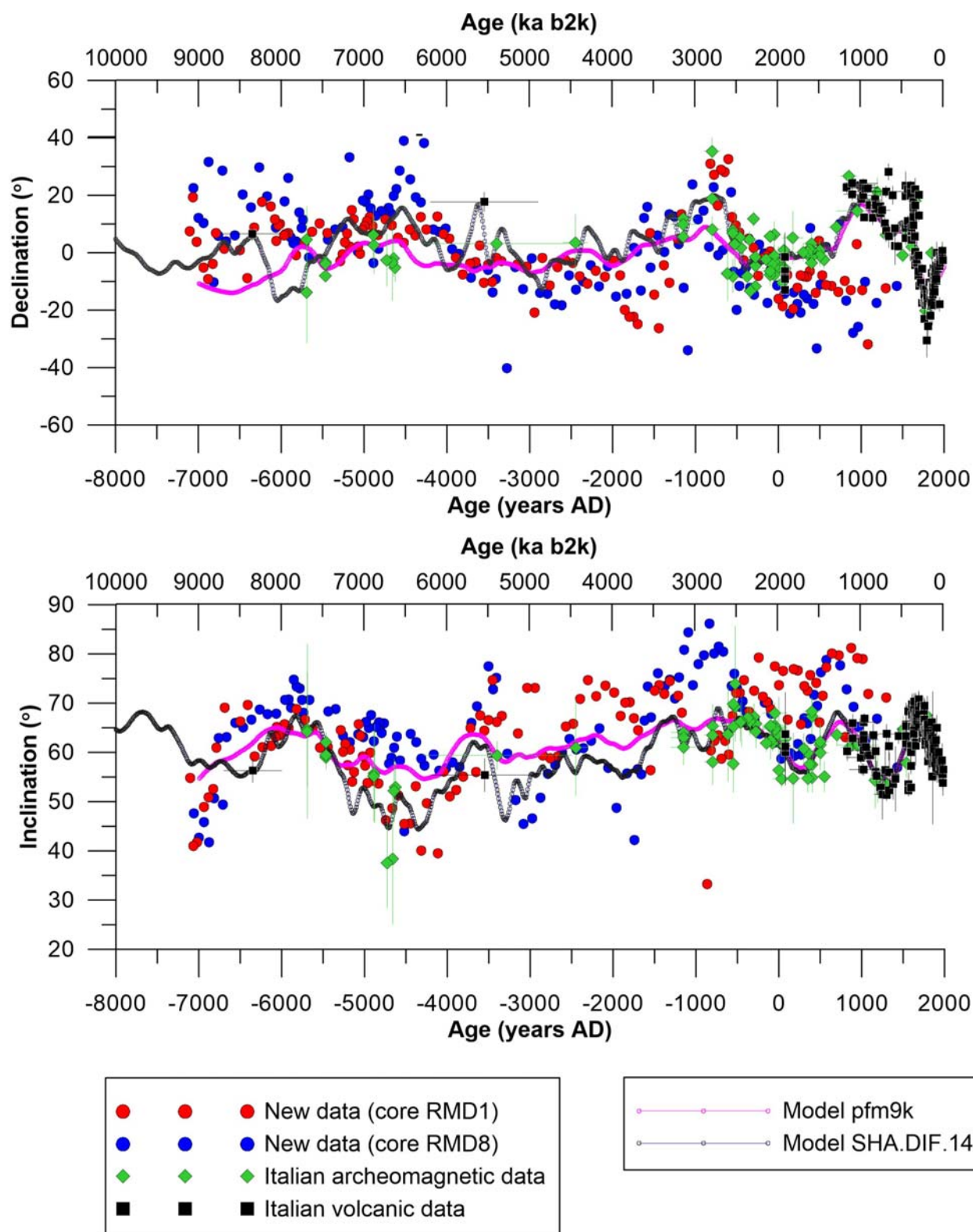
642

643 Figure 7



645 Figure 8





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649 Figure 10

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